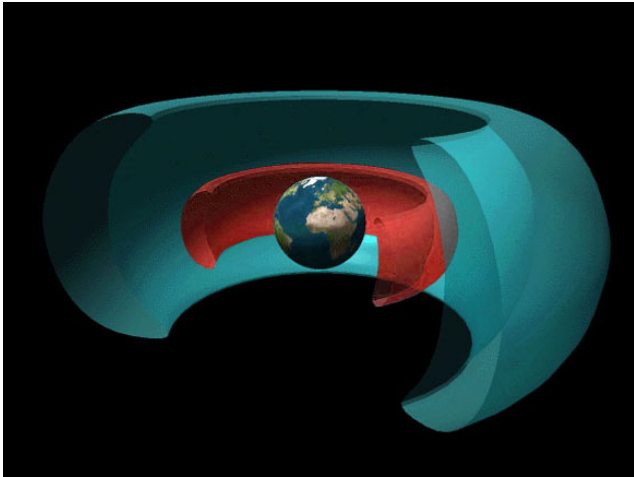
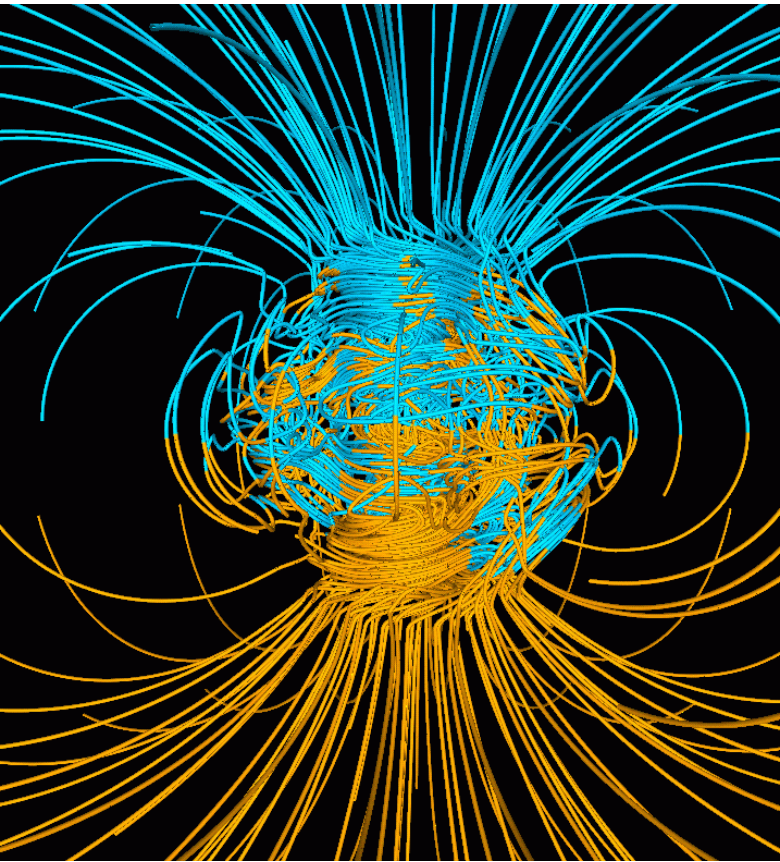
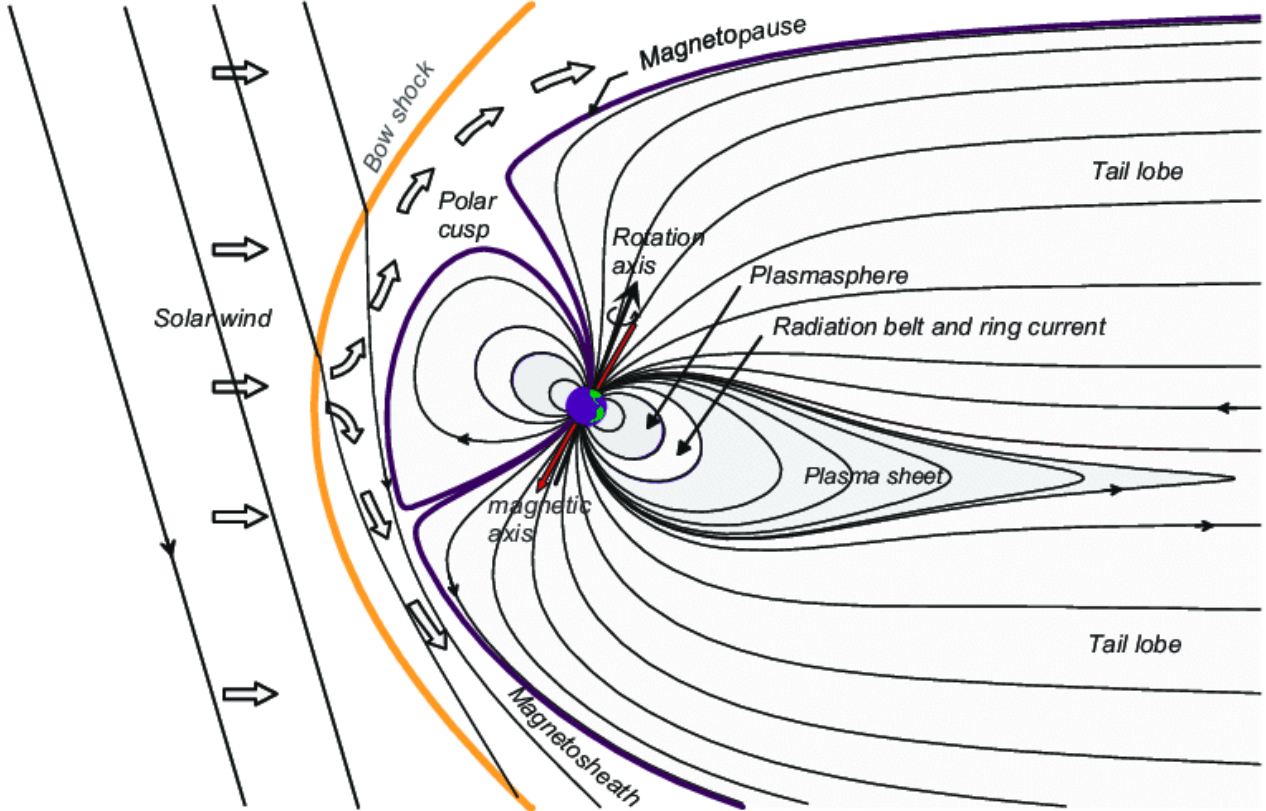


Geomagnetic field

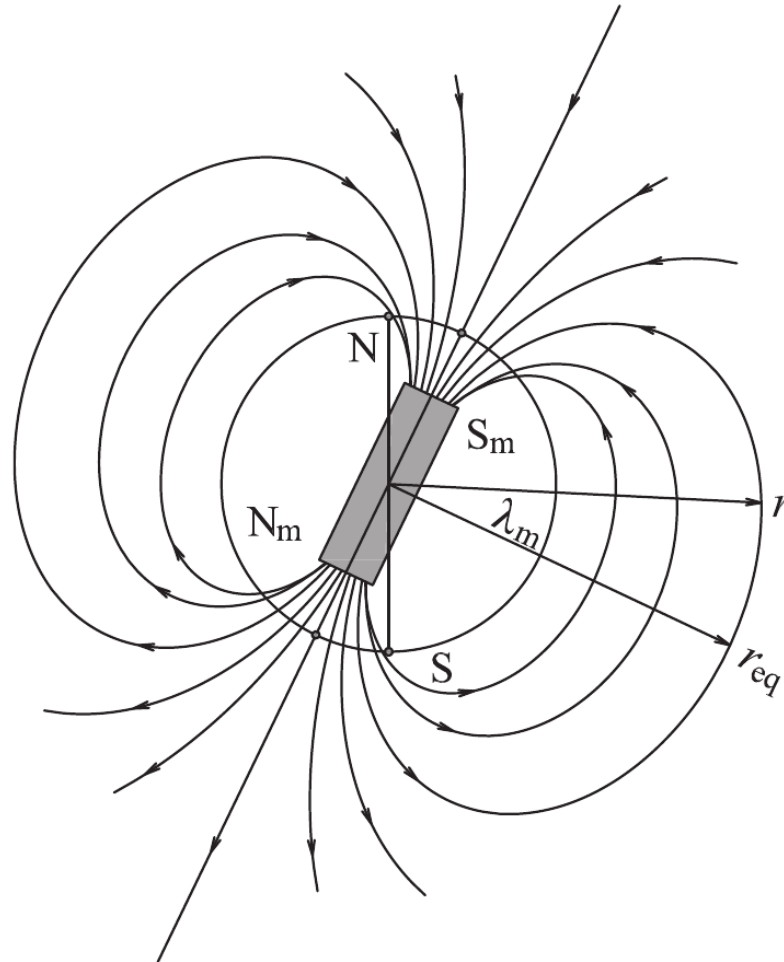


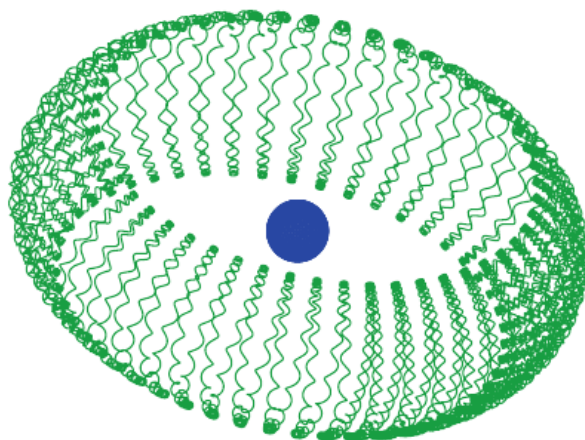
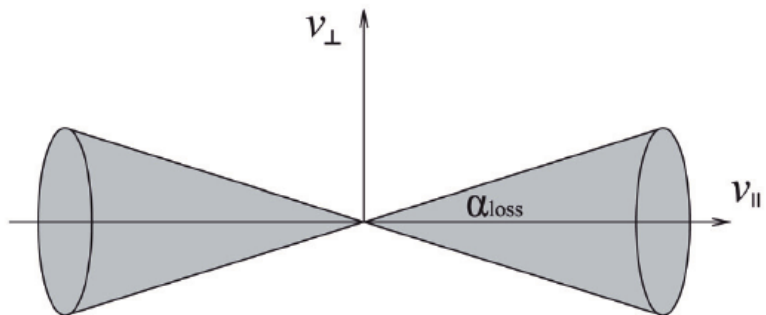
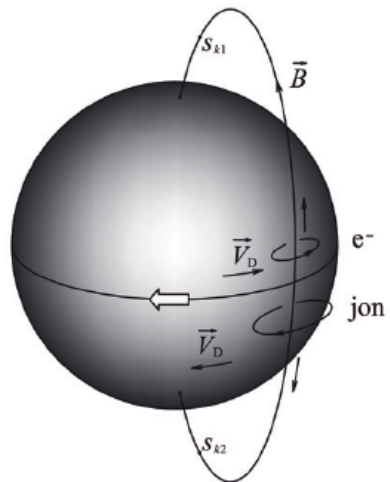
$$\vec{B} = \frac{\mu_0}{4\pi} \frac{M_{\oplus}}{r^3} (-2 \sin \lambda_m \vec{e}_r + \cos \lambda_m \vec{e}_{\lambda_m})$$

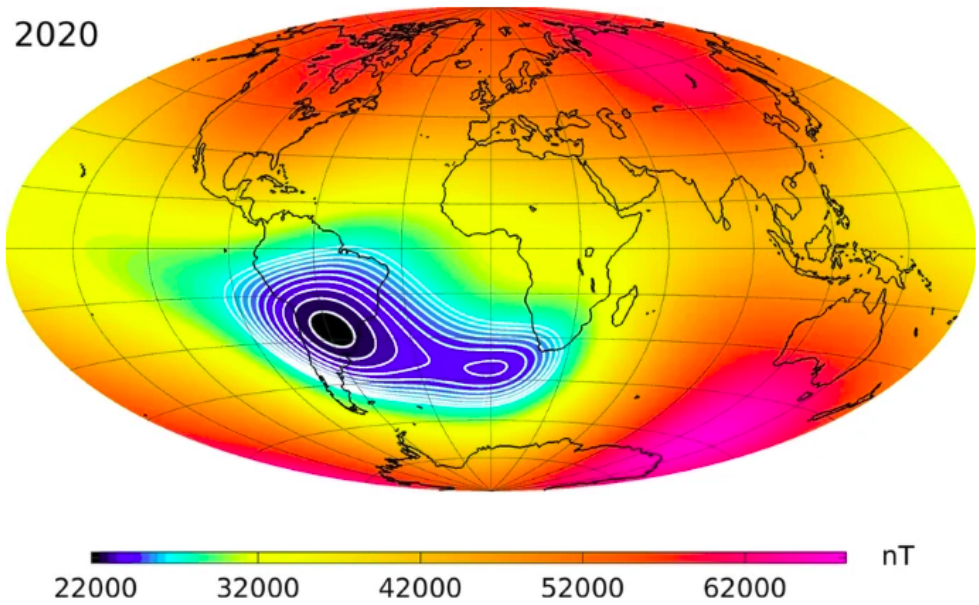
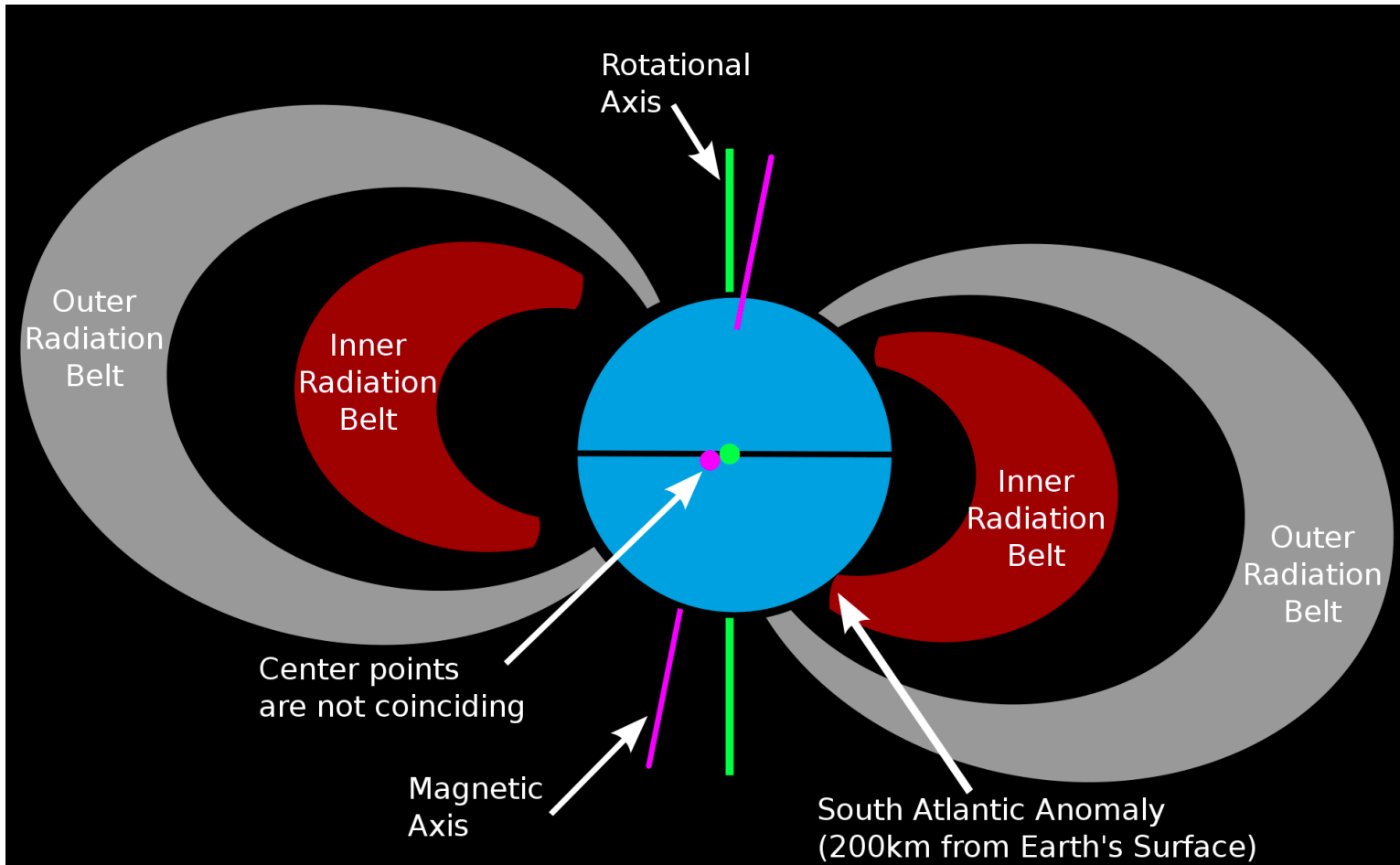
$$\mathbf{L} = r_{\text{eq}}/R_{\oplus}$$

$$M_{\oplus} \approx 8.05 \times 10^{22} \text{ A m}^2$$

$$B(\lambda_m, \mathbf{L}) = \frac{B_{\oplus}}{\mathbf{L}^3} \frac{\sqrt{1 + 3 \sin^2 \lambda_m}}{\cos^6 \lambda_m}$$

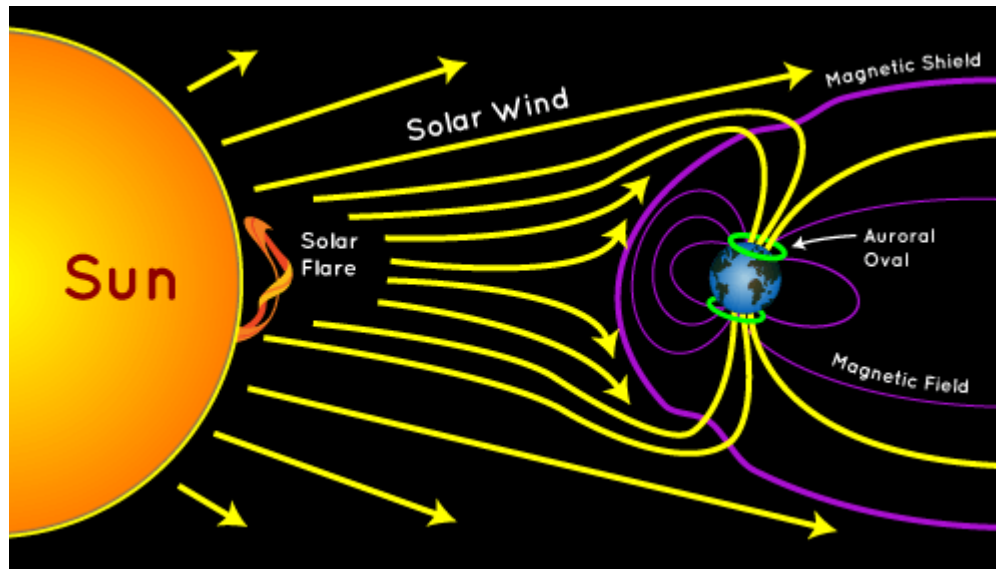






Aurora: L-values ~ 6

<https://www.youtube.com/watch?v=0I1hZCD7sT0>

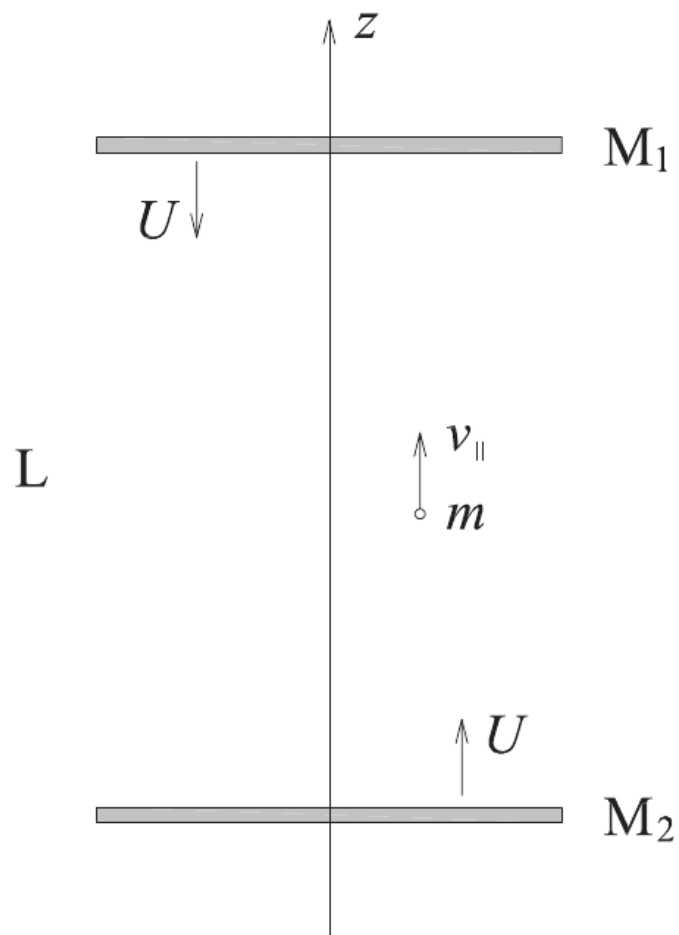


Magnetic reconnection

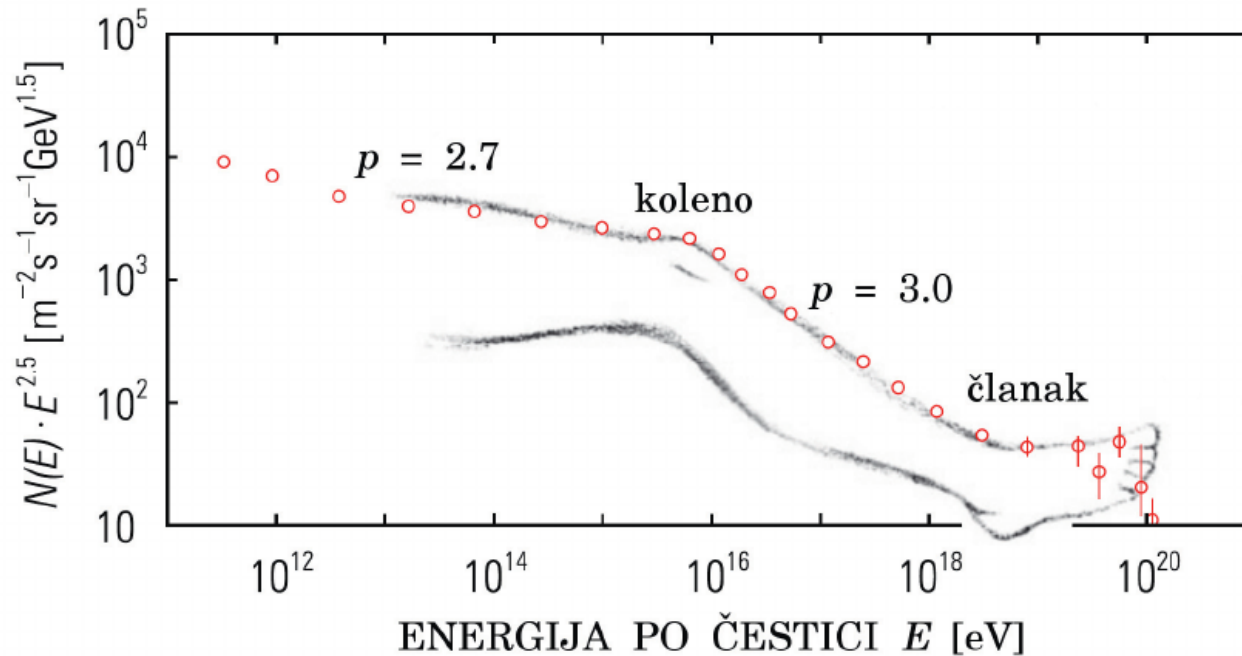
$$\oint m v_{\parallel} ds = \text{const}$$

$$p_{\parallel} L = \text{const}$$

Fermi acceleration
basic



Cosmic rays



Synchrotron radiation – high energy electrons

Pion decay – high energy ions

Cosmic rays as information carriers – problems: trajectory is changed due to magnetic field...

Diffusive Shock Acceleration in quasi-parallel shocks

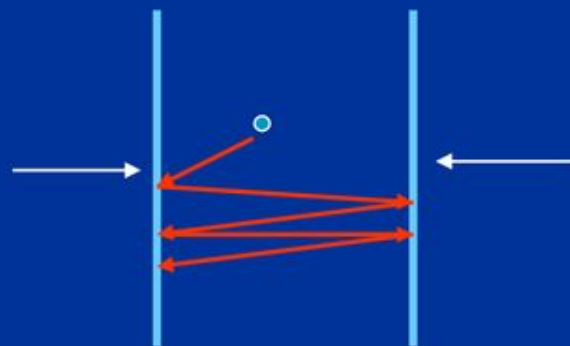
“Fermi first order process”

Alfvén waves in a converging flow
act as converging mirrors

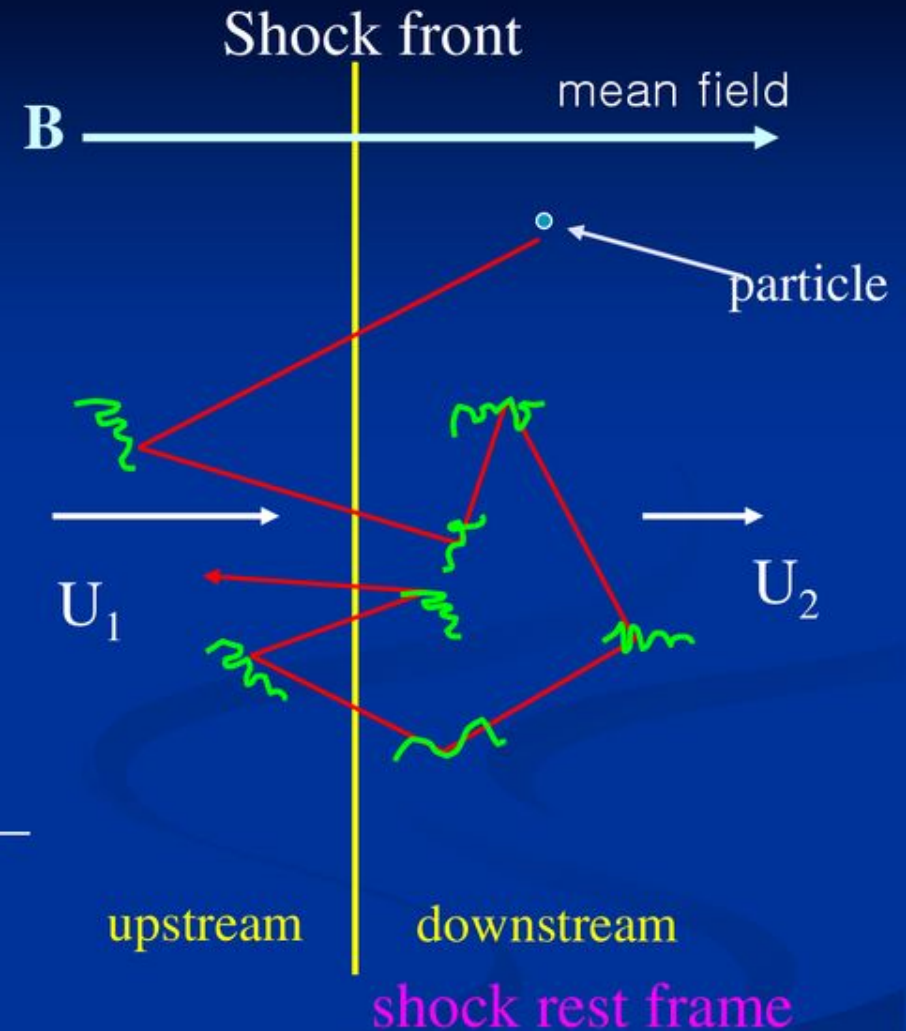
- particles are scattered by waves
- cross the shock many times

$$\frac{\Delta p}{p} \sim \frac{U_s}{v} \quad \text{energy gain at each crossing}$$

$$v \approx c$$



Converging mirrors



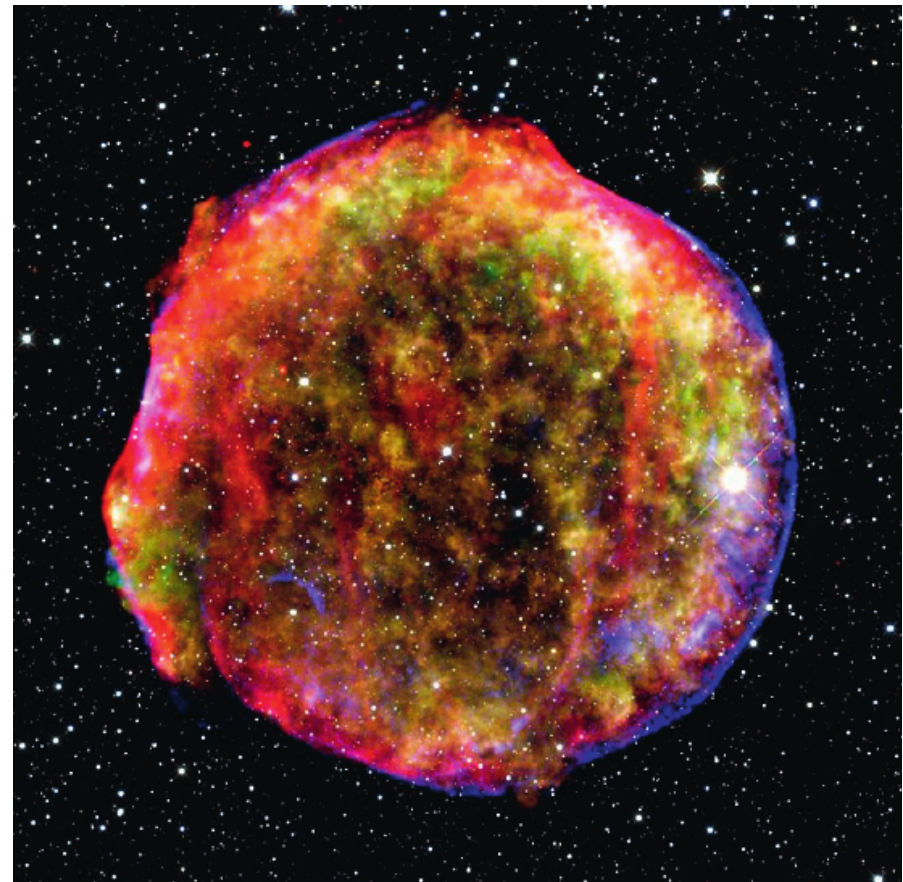
Incoherent continuum radiation can be treated in orbital method

Calculate the emissivity of one particle – find distribution of particles $N(E)$ – calculate how system of particles radiate

Bremsstrahlung, synchrotron, inverse Compton, ...

Thermal vs. non-thermal radiation

$$I = \frac{1}{4\pi\epsilon_0} \frac{2}{3} \frac{q^2}{c^3} \dot{v}^2$$



Cyclotron resonance

wave-particle interaction

$$\vec{B} = B\vec{e}_z, \quad \vec{E} = E \cos(\omega t)\vec{e}_x$$

$$\ddot{x} = \omega_c \dot{y} + \frac{q}{m} E \cos(\omega t), \quad \ddot{y} = -\omega_c \dot{x}, \quad \ddot{z} = 0$$

if $\omega = \omega_c$

$$r_c \equiv v_{\perp}^0 / \omega_c$$
$$x = r_c(1 - \cos(\omega_c t)) + \frac{q}{m} \frac{Et \sin(\omega_c t)}{2\omega_c},$$
$$y = r_c \sin(\omega_c t) + \frac{q}{m} \frac{E}{2\omega_c^2} (\omega_c t \cos(\omega_c t) - \sin(\omega_c t))$$